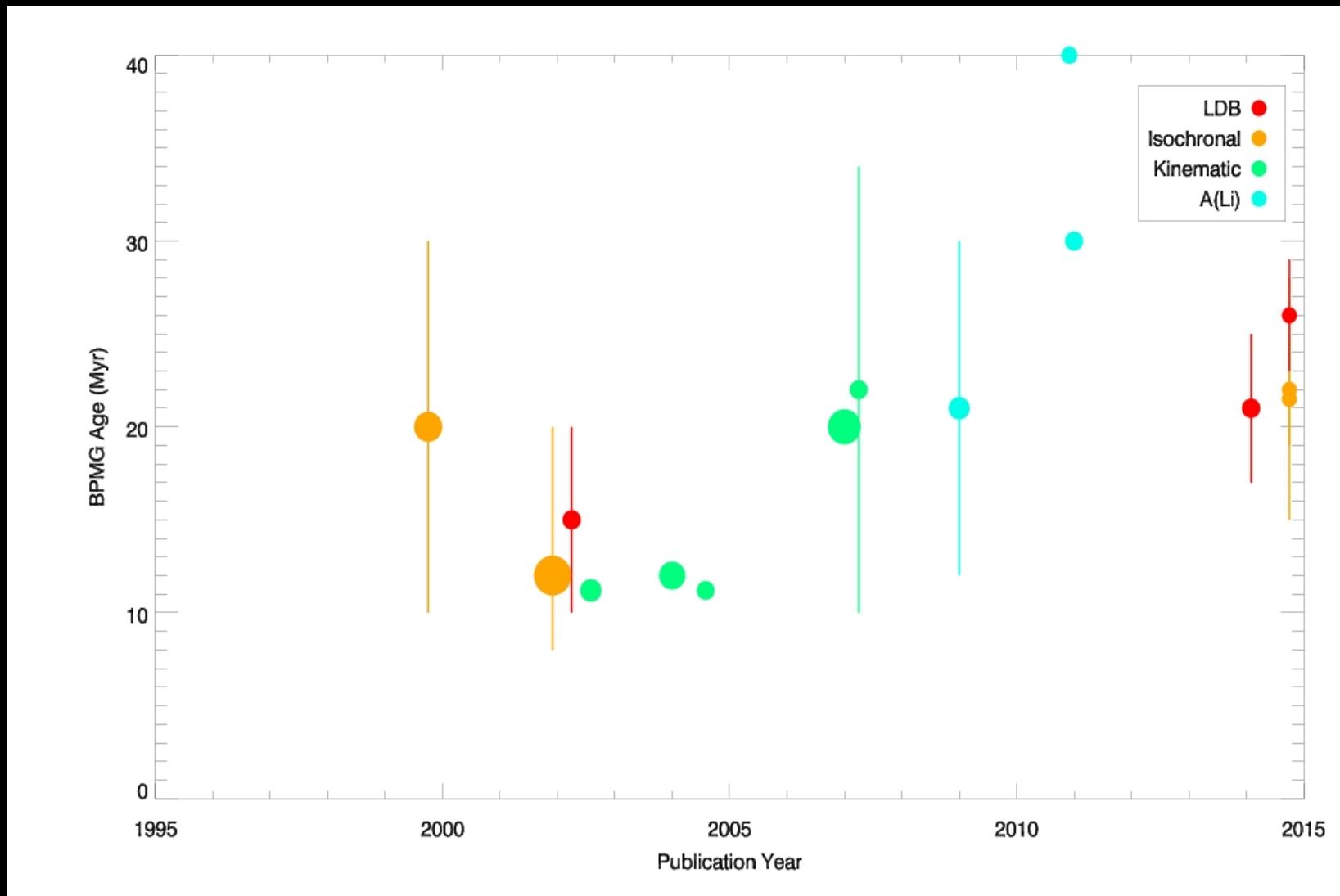


How old is Beta Pic?

Alex Binks, Keele University
Paris, 8th September 2014

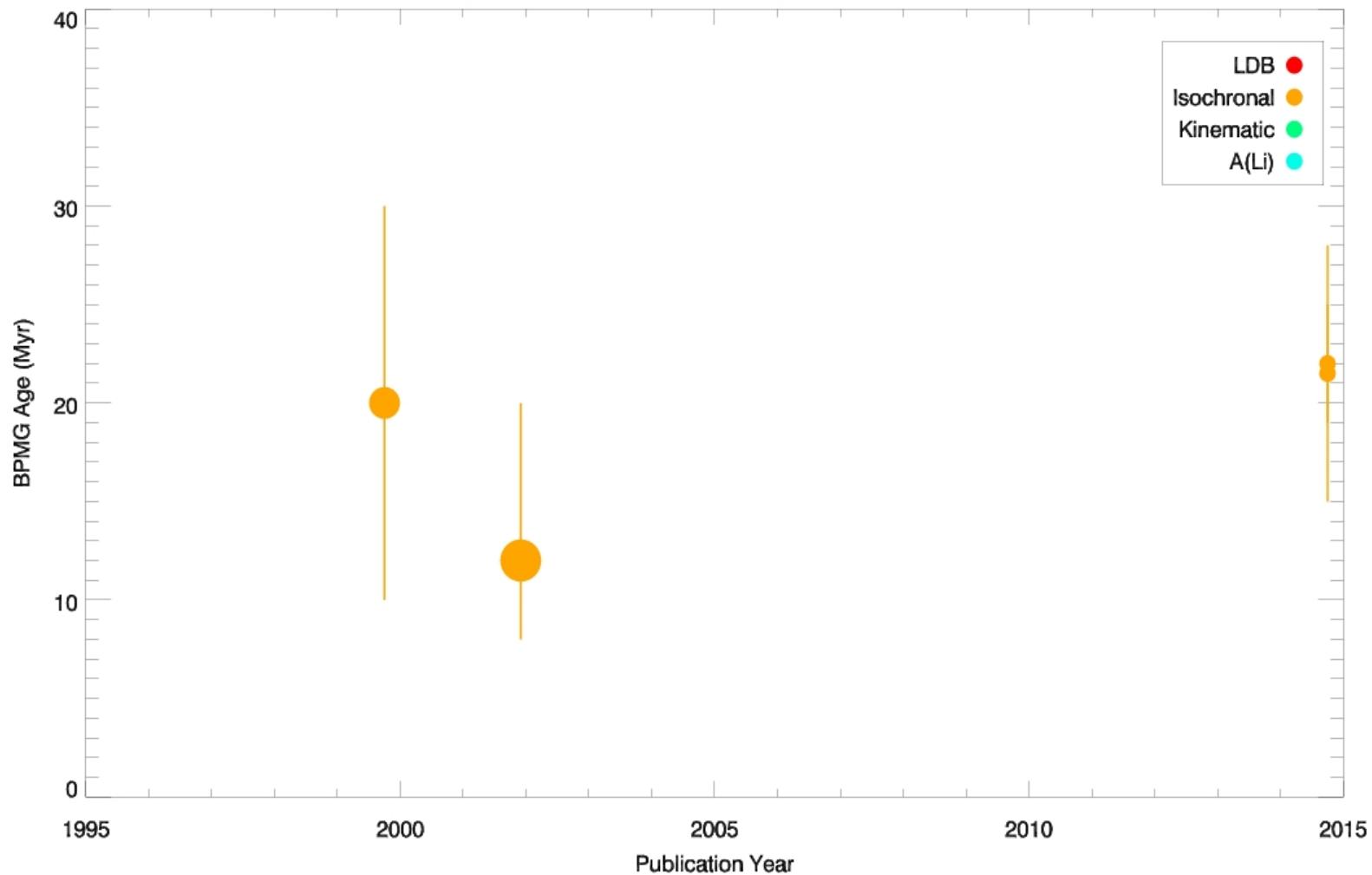
The age of BPMG over time



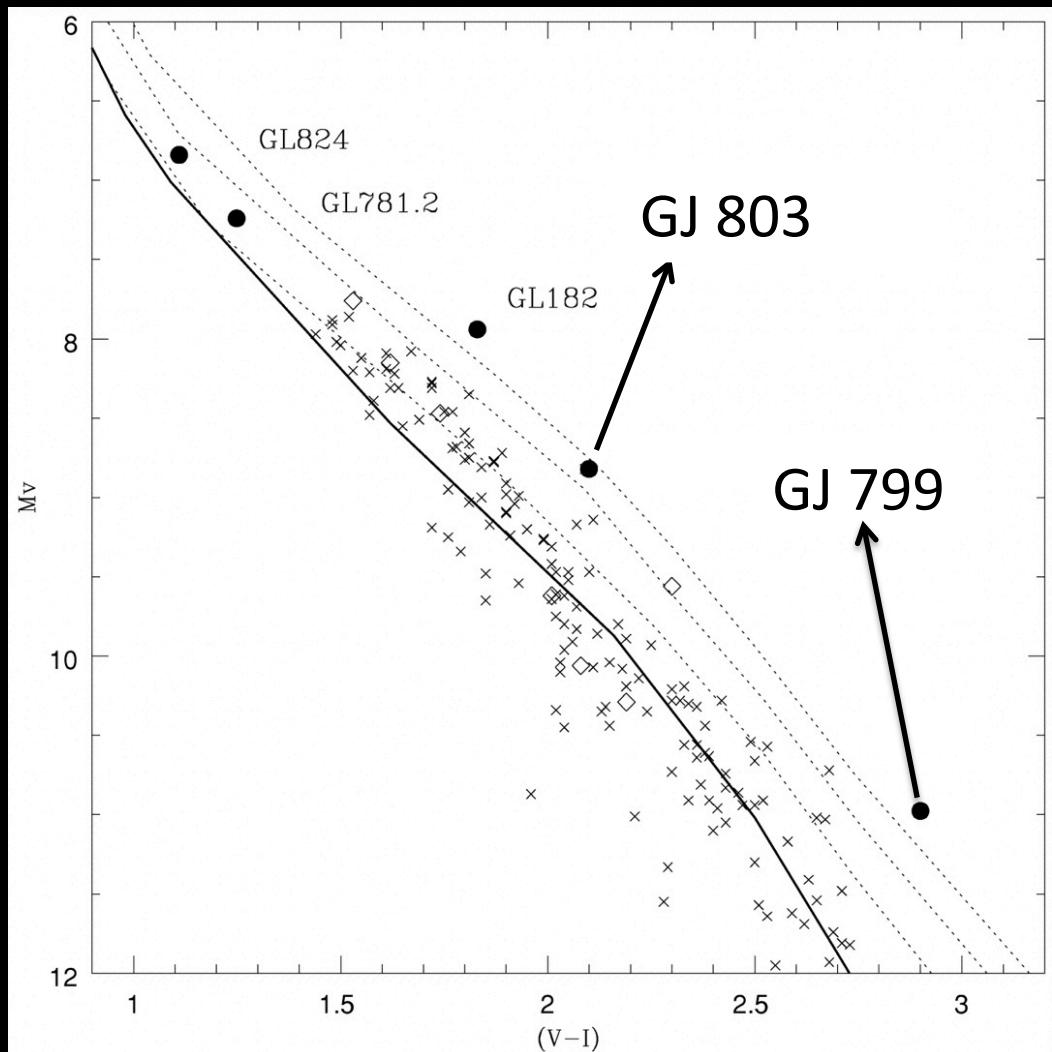
Early age estimates

AGE (Myr)	REASON	REFERENCE
≤ 200	Under-luminous compared to MS.	Paresce 1991
As young as 10	Debris disk compared to young analog HR4796.	Jura et al. 1993
Double valued, 12 or 3000	HR diagram, debris disk.	Lanz et al. 1995
≥ 100	Sublimating comets can replenish disk material.	Brunini et al. 1996
8 – 100	Near to, or just on the ZAMS.	Crifo et al. 1997
200 ± 100	POTENTIALLY PART OF A MOVING GROUP.	Barrado 1998

Isochronal ages



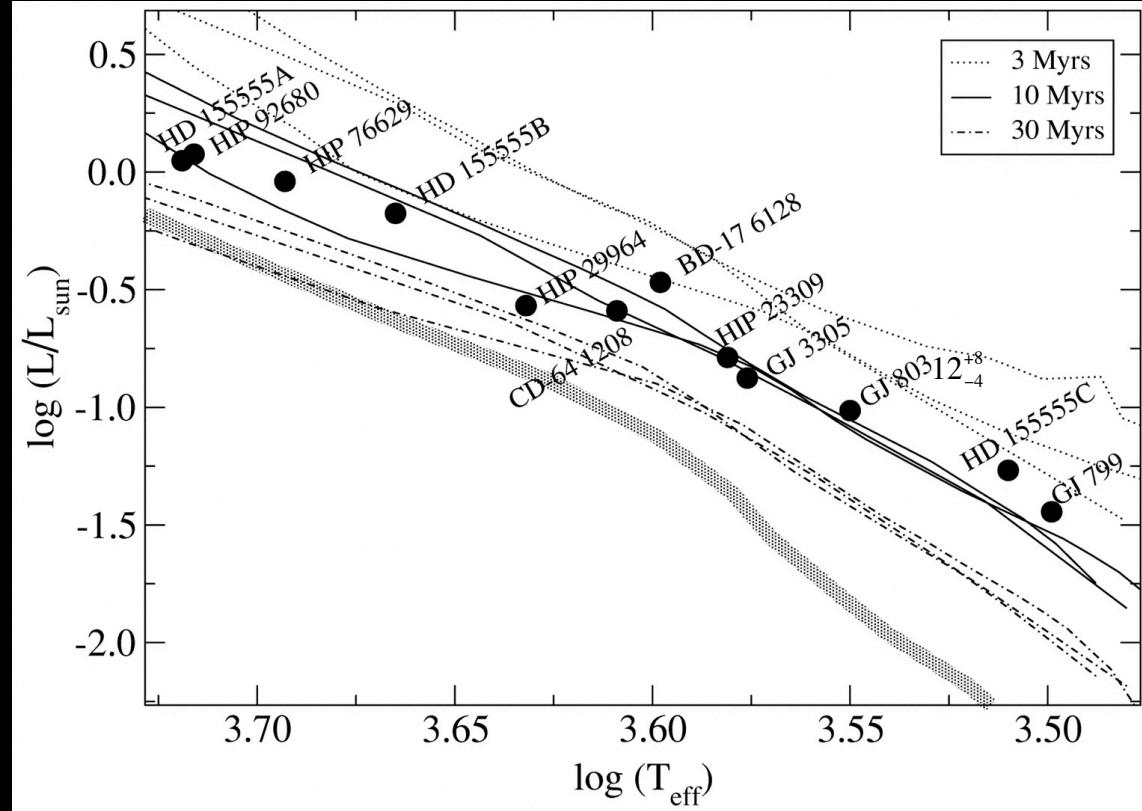
Barrado y Navascues et al. 1999



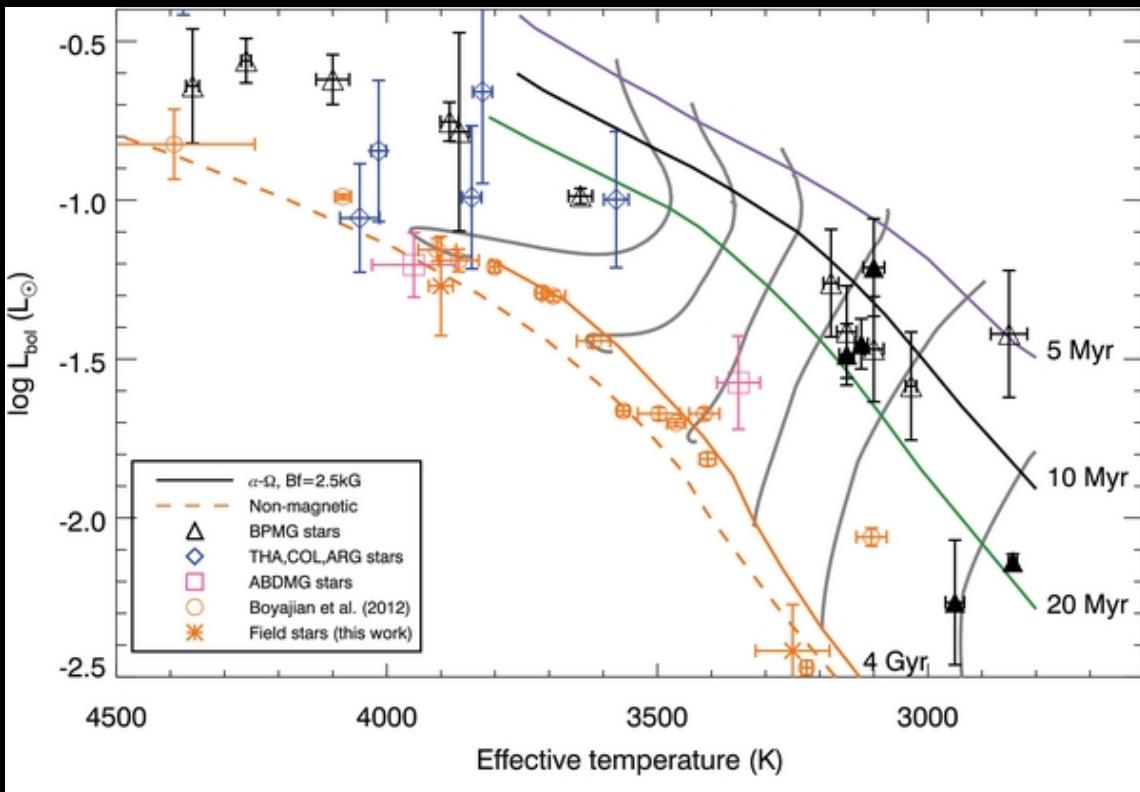
- CMD (KM stars, M_V vs $V-I$)
- 2 objects: GJ799, GJ803
The BPMG!
- UVW and age
- Age: 20 ± 10 Myr

Zuckerman et al. 2001

- HR diagram (GKM)
- 18 new star systems
- Convert M_V vs $V-I$ to theoretical plane
- Age: 12^{+8}_{-4} Myr

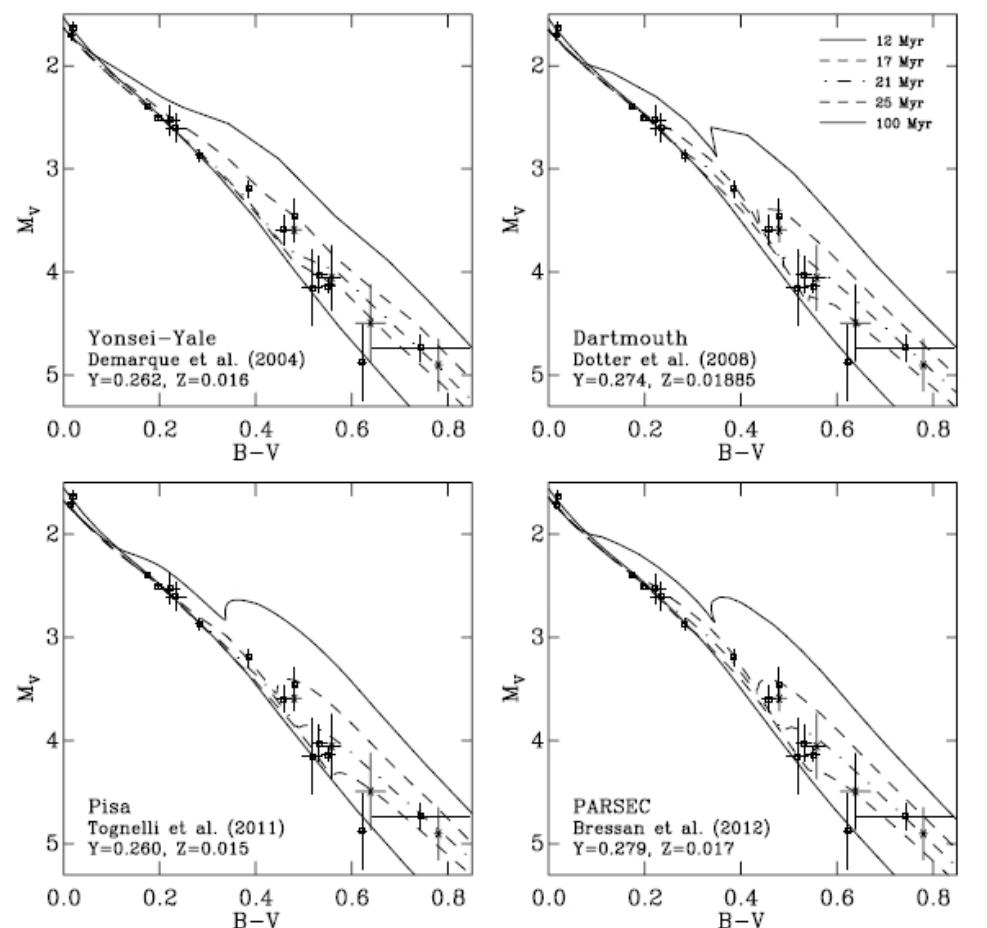


Malo et al. 2014



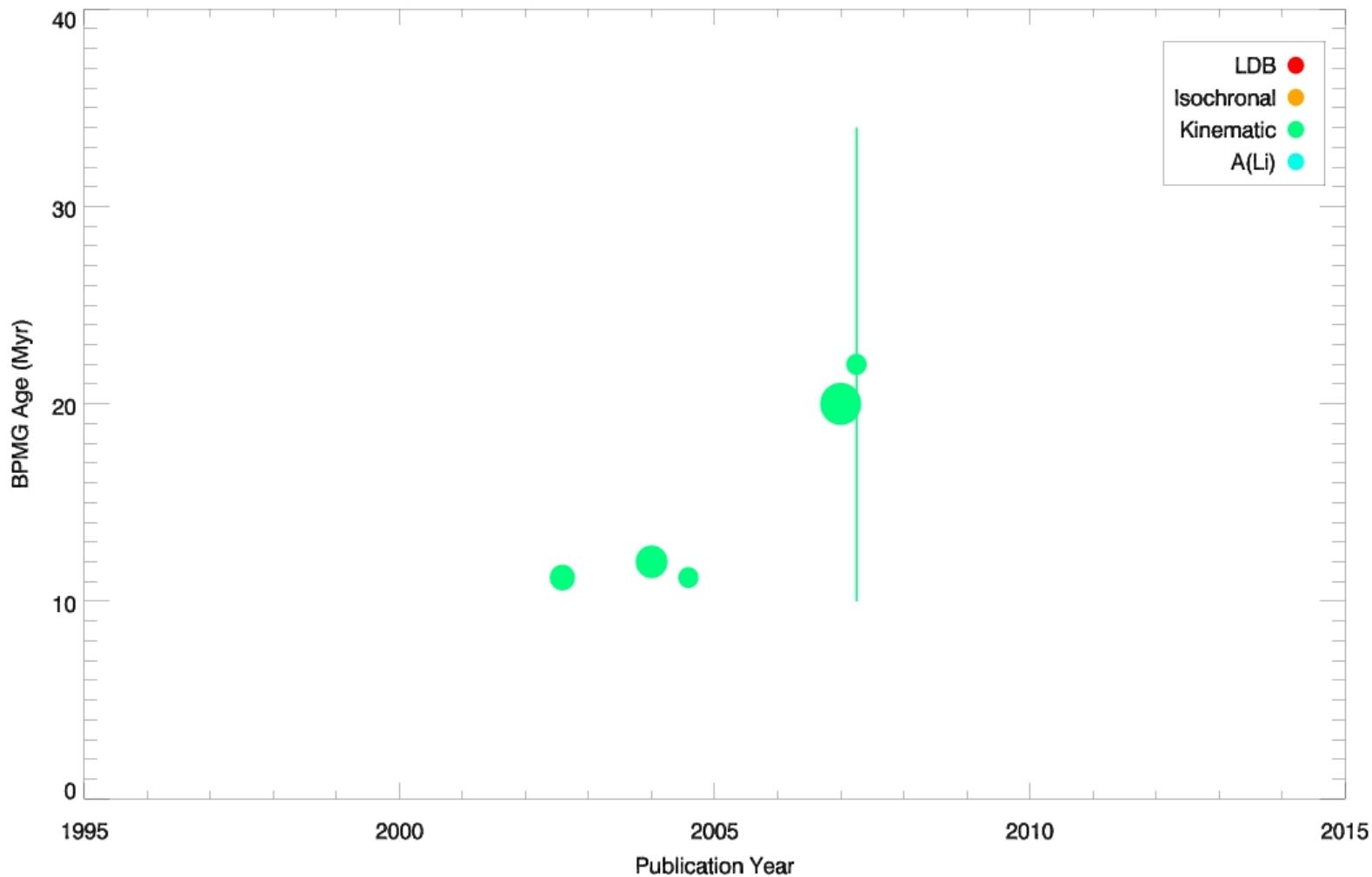
- HR diagram (KM)
- Sample of ‘bonafide’ stars from Malo et al. 2013
- Dartmouth models, $\langle B \rangle = 2.5 \text{ kG}$
- Age: $21 \pm 6.5 \text{ Myr}$

Mamajek & Bell 2014 (in prep)

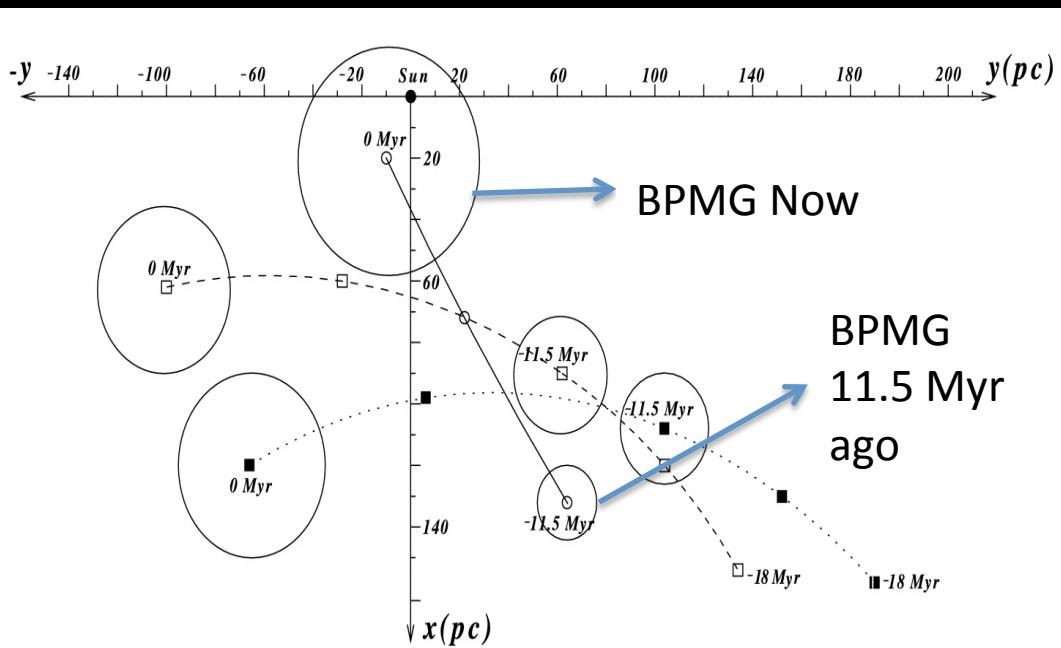


- CMD (AFG stars, M_V vs $V-I$)
- Revised astrometry
- 4 separate evolutionary models
- Age: 22 ± 3 Myr

Kinematic ages

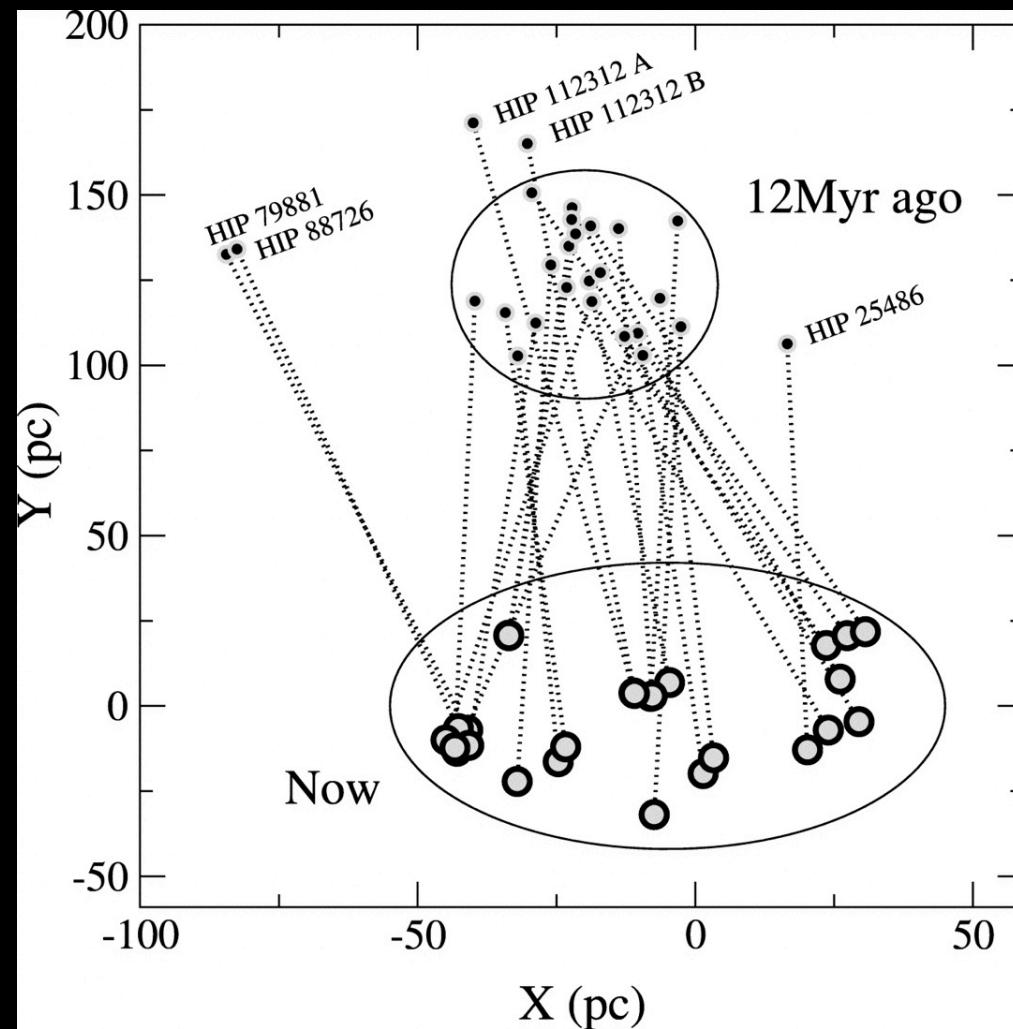


Ortega et al. 2002, 2004



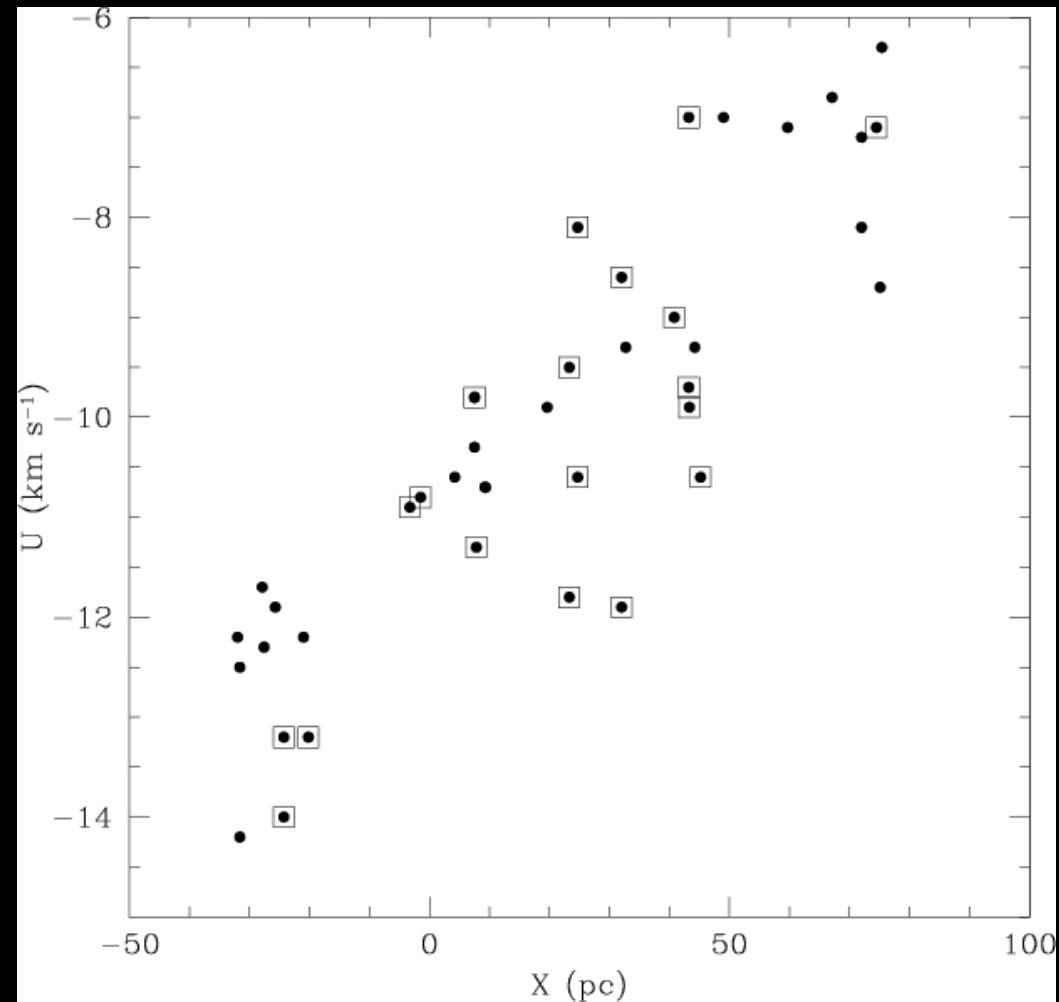
- Traceback orbits of all BPMG members
- First min at -11.5 Myr (1/3 of today's size)
- Birthsite near UCL, LCC?
- Error bars?

Song et al. 2003



- ‘Only works < 15 Myr’
- Observe a min at 12 Myr
- Positional error of 12 pc based on HIP112312AB
- Outliers?

Torres et al. 2006

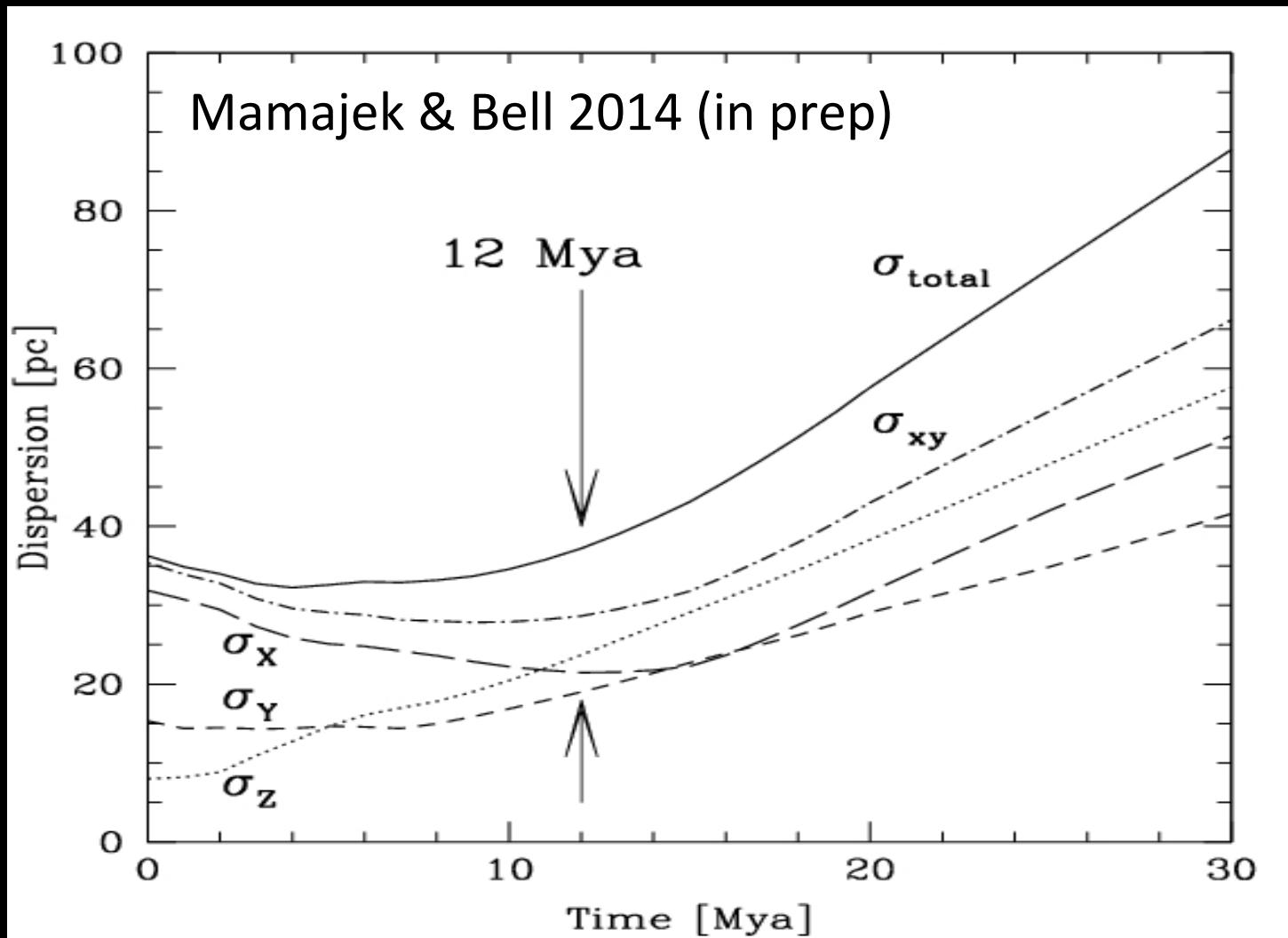


- Observe how the group has expanded
- $dU/dX = 0.053 \text{ km s}^{-1} \text{pc}^{-1}$
- Age: 18 Myr
- No error analysis

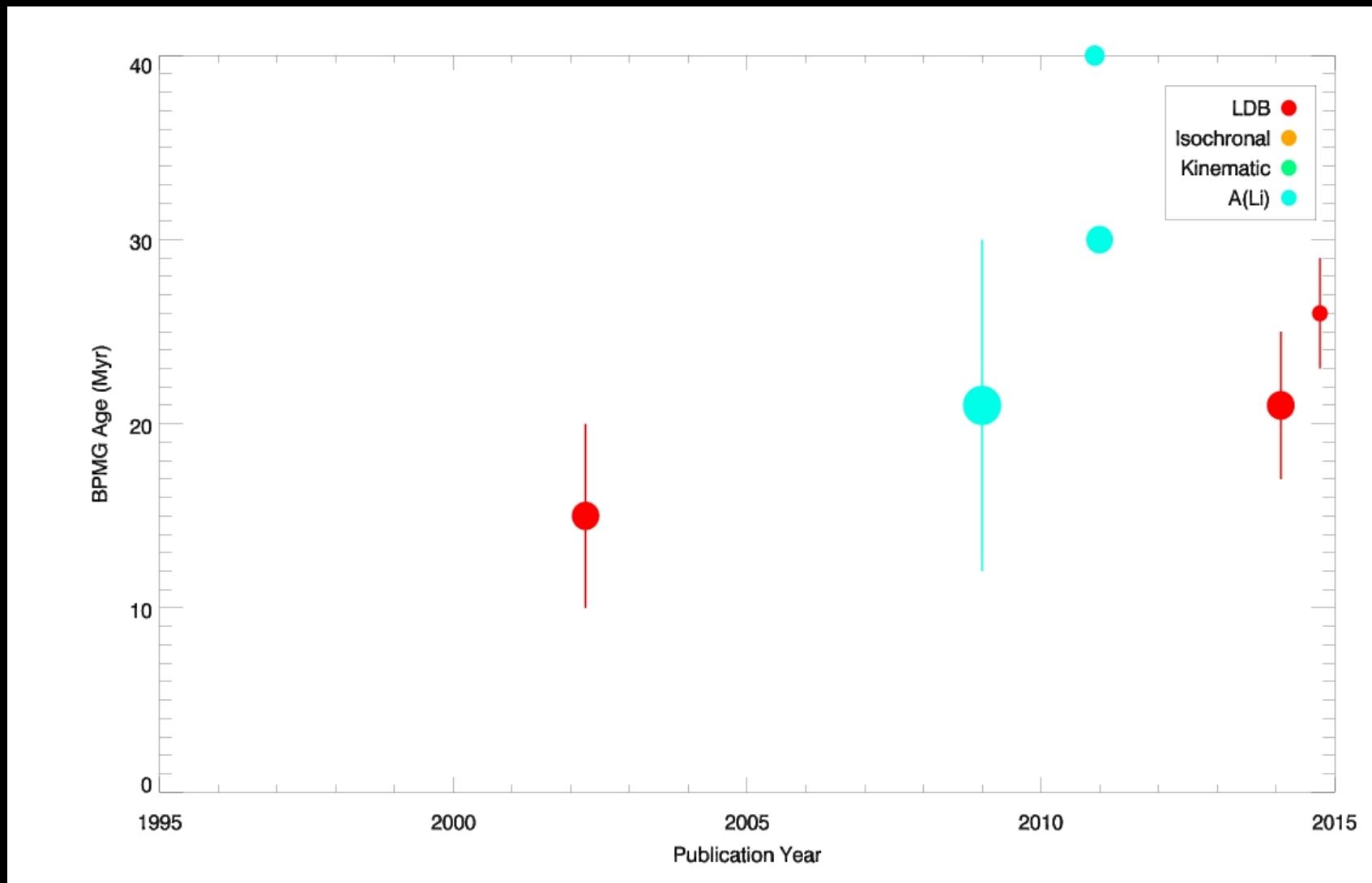
Makarov 2007

- Track the times of ‘closest encounter` for 14 members
- Excludes β Pic (!) and HIP 29964
- Mean time of nearest approach: 22 ± 12 Myr
- Weak expansion and larger scatter

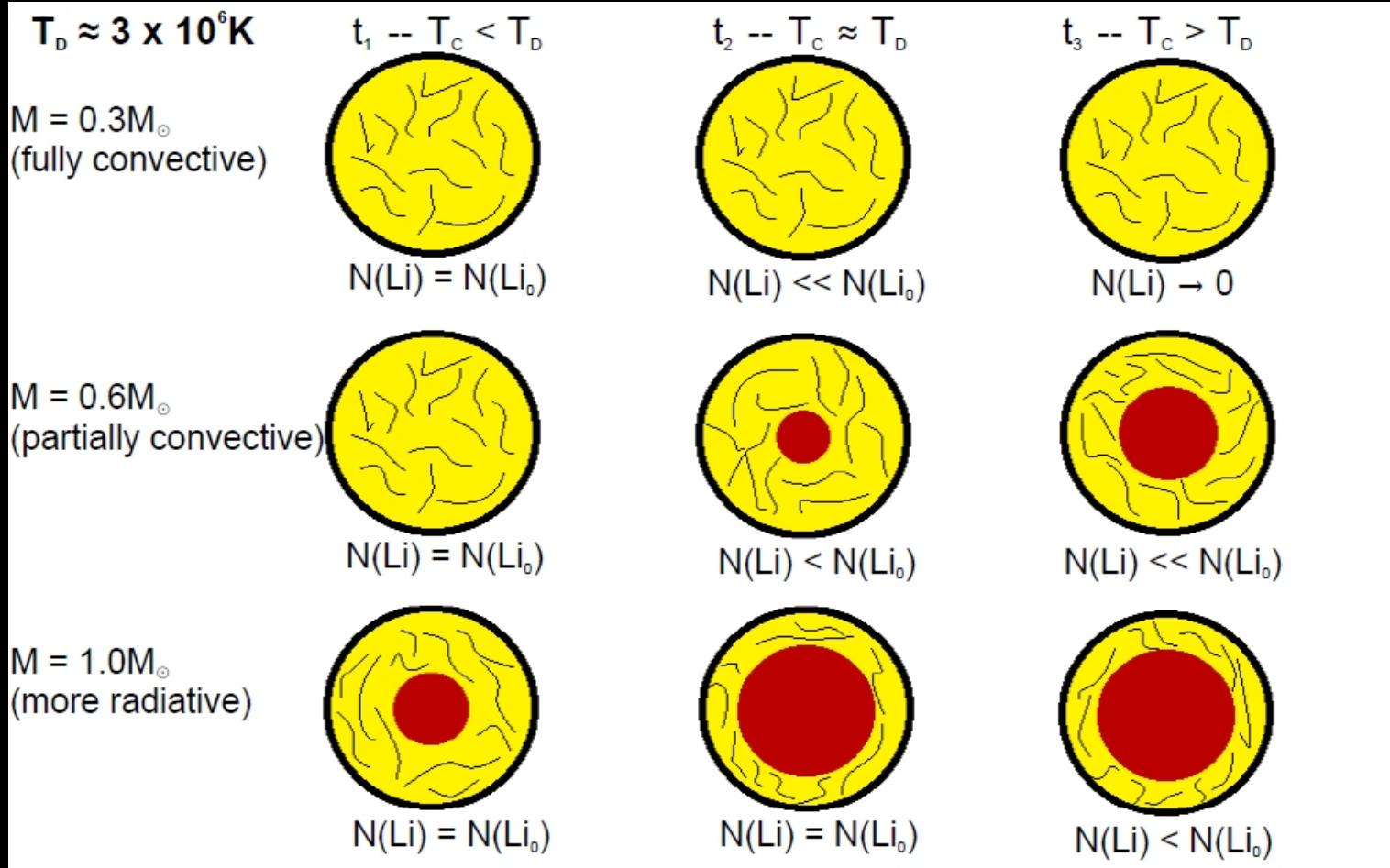
In light of new results...



LDB (and Li depletion) ages



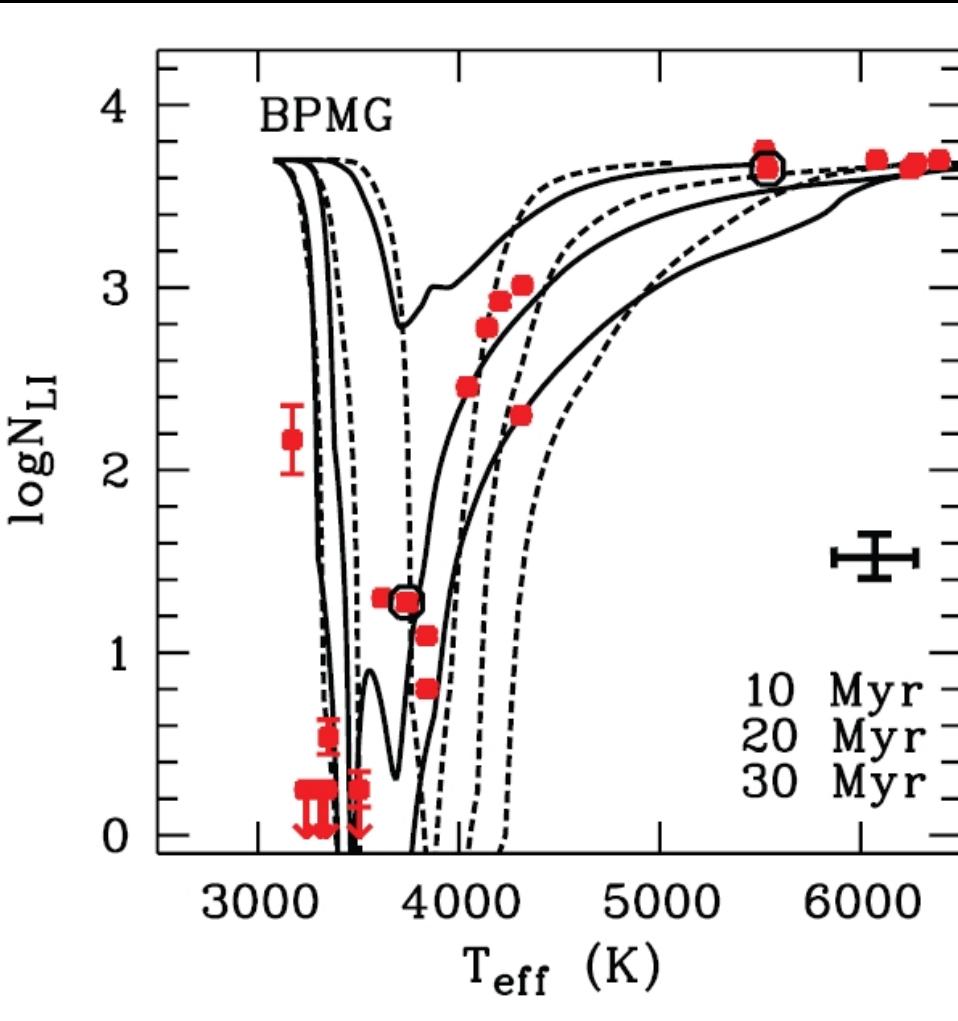
LDB technique



Song et al. 2002

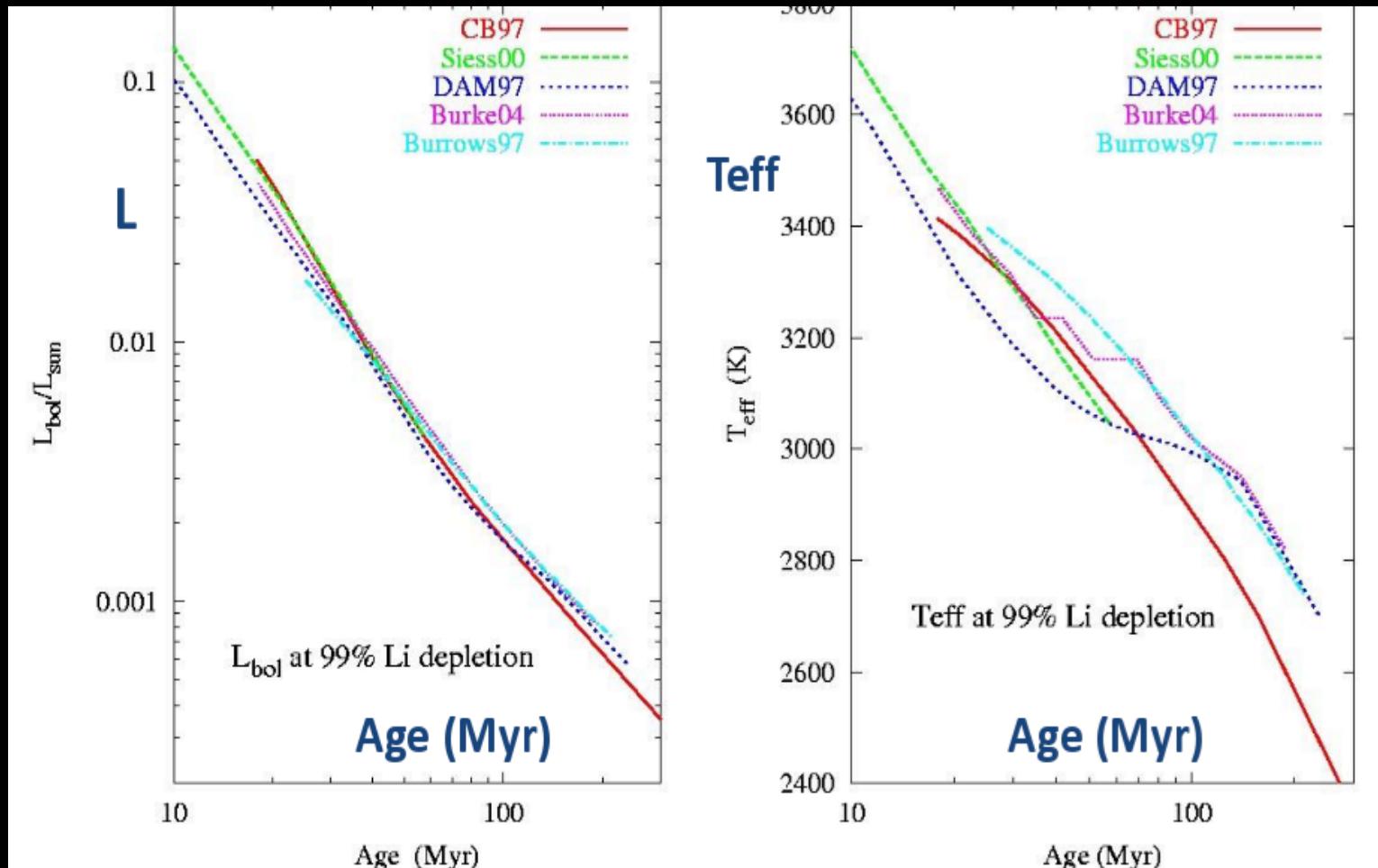
- Visual binary system HIP112312AB
- Li-rich object (M4.5)
- Li-poor object (M4.0)
- LDB age: 15 ± 5 Myr

Li Depletion: Mentuch et al. 2008 and Yee & Jensen 2010

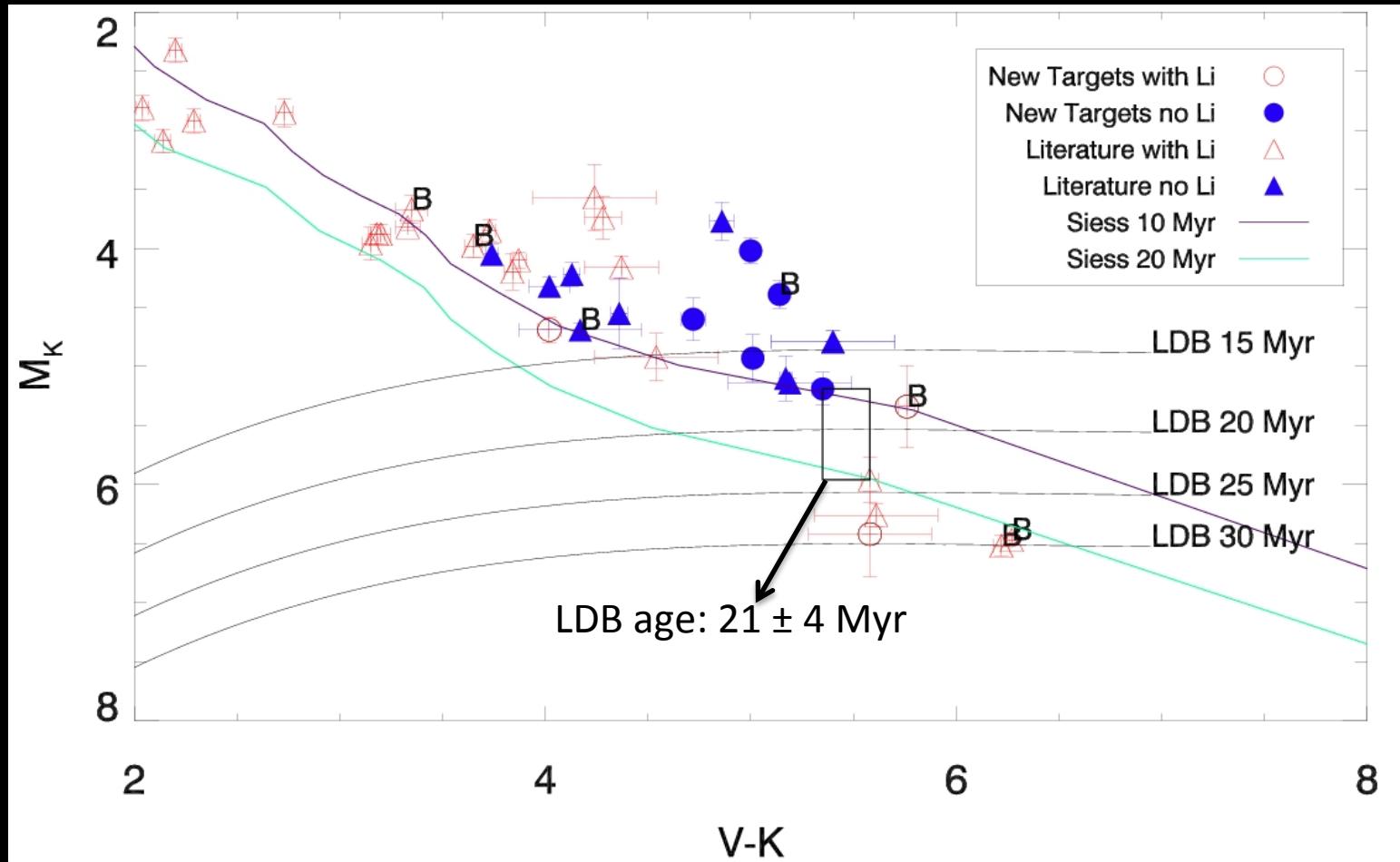


- Both use FGKM stars
- Both use T_{eff} in the models
- Mentuch: 21 ± 9 Myr
- Yee: ≈ 30 Myr

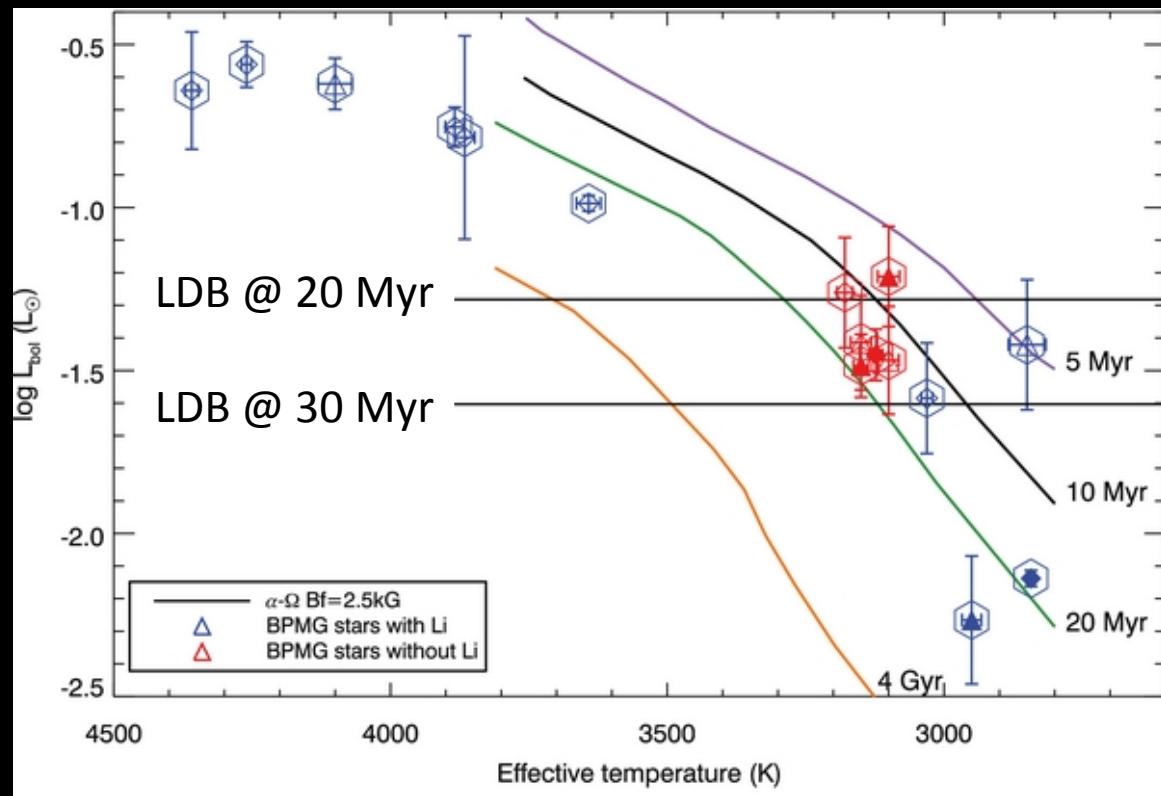
Model dependency of the LDB



Binks & Jeffries 2014



Malo et al. 2014

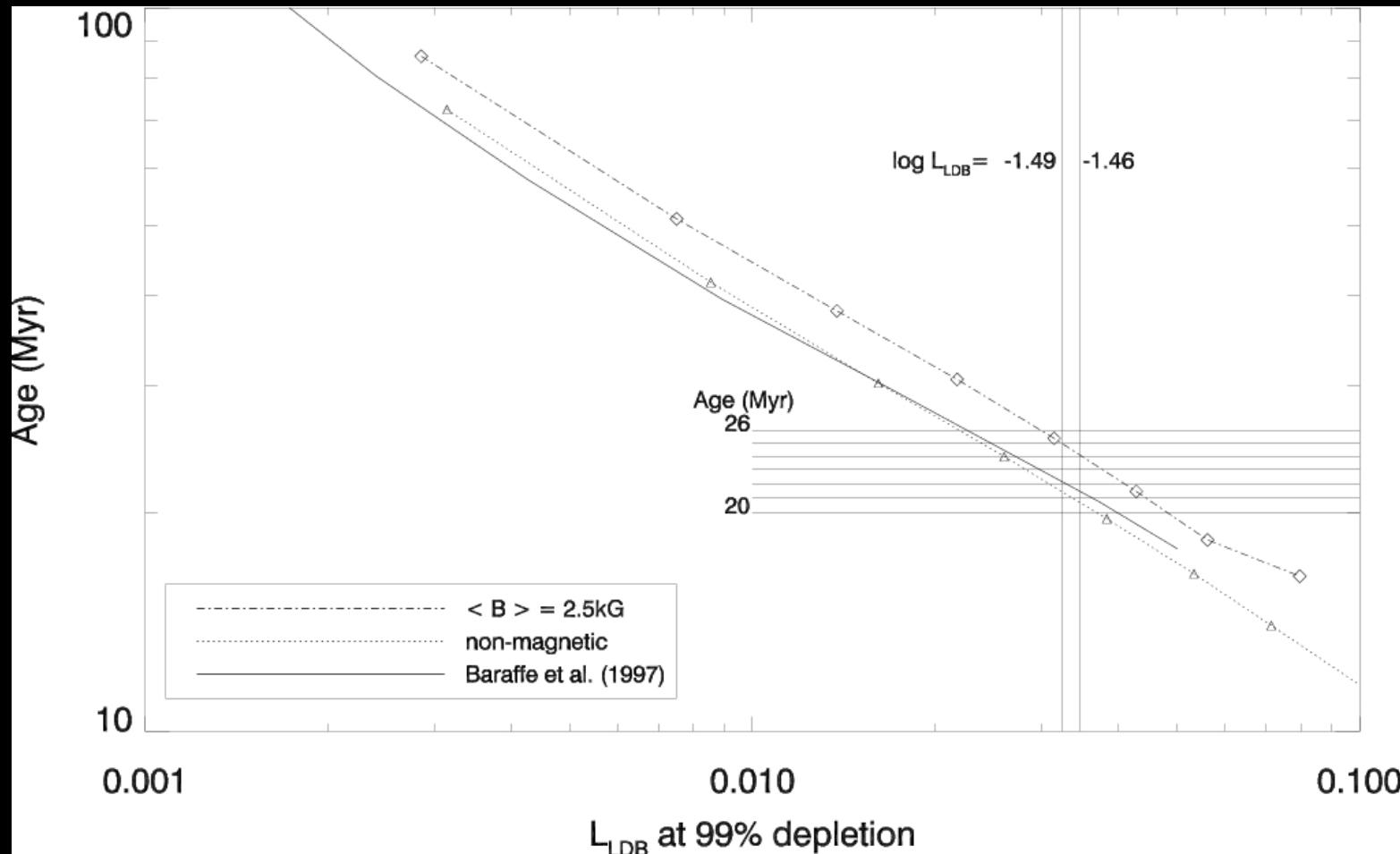


- Identify LDB at 99% depletion
 - Use magnetic Dartmouth models
 - LDB age: 26 ± 3 Myr

Binks & Jeffries 2014:
 21 ± 4 Myr, $\log L = -1.46$

LDB comparisons

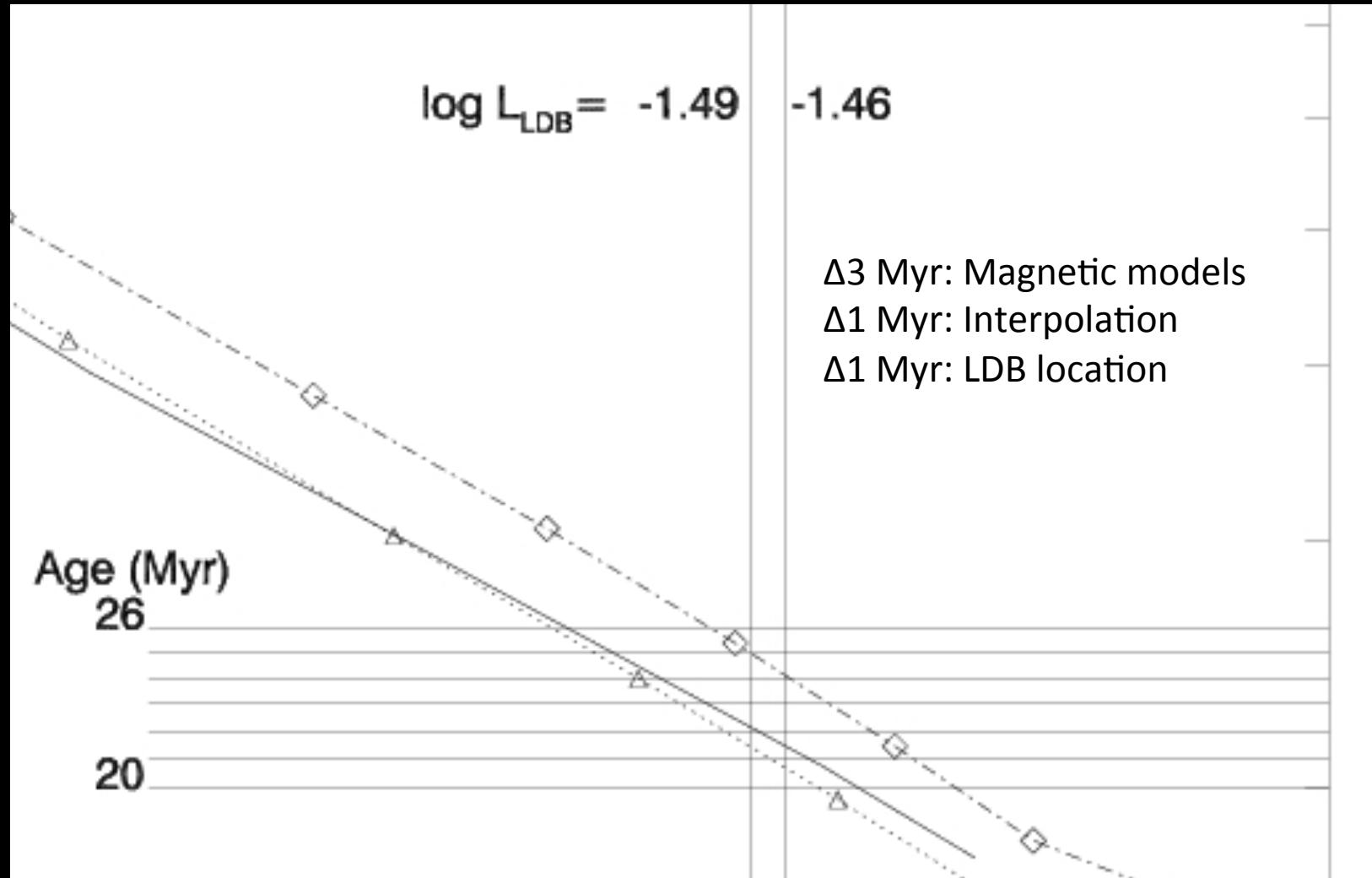
Malo et al. 2014:
 26 ± 3 Myr, $\log L = -1.49$



Binks & Jeffries 2014:
 21 ± 4 Myr, $\log L = -1.46$

LDB comparisons

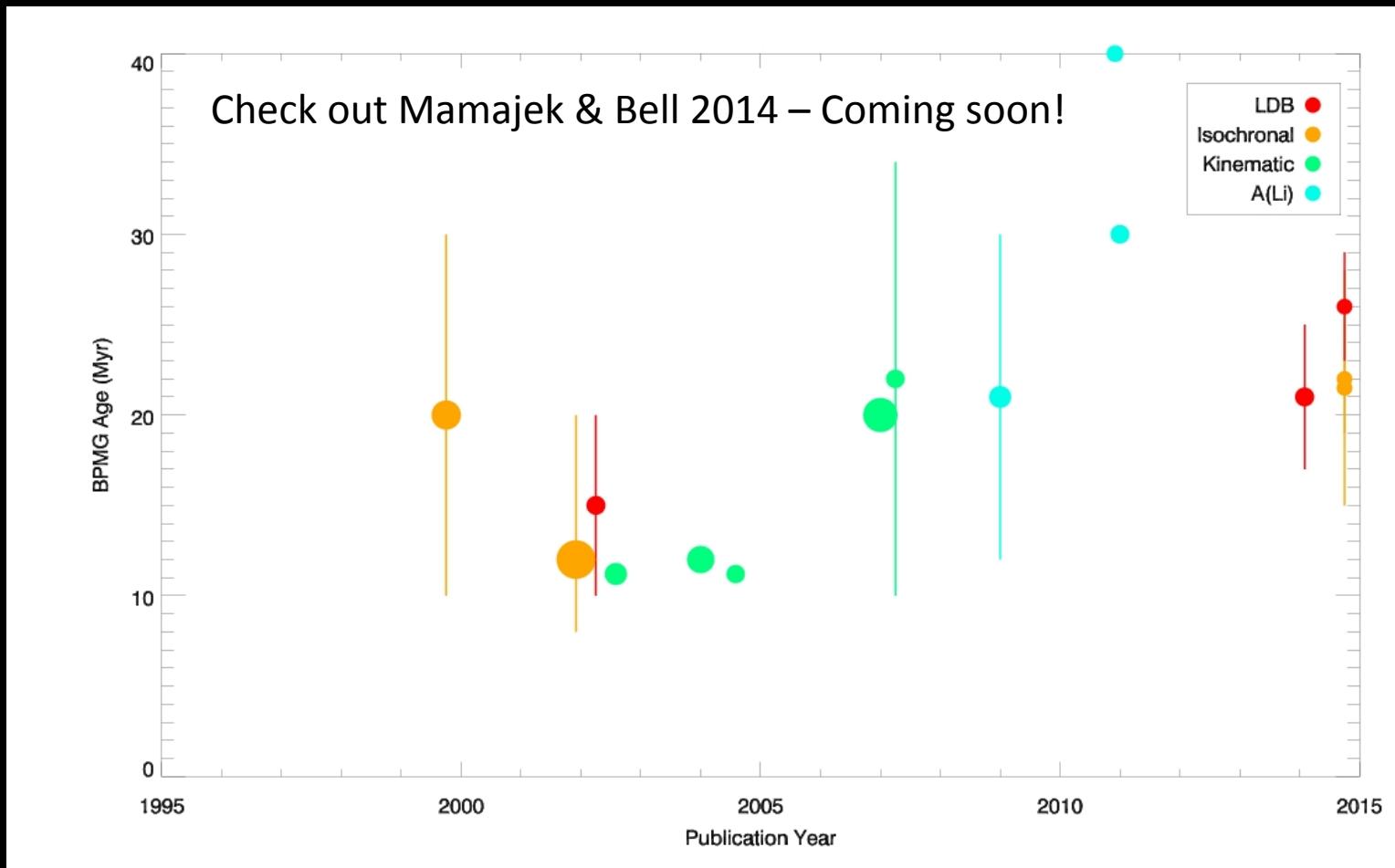
Malo et al. 2014:
 26 ± 3 Myr, $\log L = -1.49$



So how old is β Pic?

- Recent isochronal ages suggest β Pic is about double the Zuckerman et al 2001 age
- Kinematic ages are ‘model-independent’, but potentially hazardous due to astrometry errors and sample choice – GAIA?
- LDB suggests an age of 20 – 25 Myr
- Combining all age estimates: 23 ± 4 Myr

Thanks for listening!



References

- Paresce (1991), A&A, 247, 1
Jura et al. (1993), ApJL, 418, L37
Lanz et al. (1995), ApJL, 447, L41
Brunini et al. (1996), MNRAS, 283, L84-L88
Crifo et al. (1997), Conf Prod, 437-440
Barrado (1998), Astrophysics and Space Science, 263, 235
Barrado et al. (1999), ApJ, 520, L123-L126
Zuckerman et al. (2001), ApJ, 562, L87-L90
Ortega et al. (2002), ApJ, 575, 75
Song et al. (2002), ApJ, 581, 43
Song et al. (2003), ApJ, 599, 342
Ortega et al. (2004), ApJ, 609, 243
Torres et al. (2006), A&A, 460, 695
Makarov et al. (2007), ApJ, 658, 480
Mentuch et al. (2008), ApJ, 689, 1127
Yee & Jensen (2010), ApJ, 711, 303
Malo et al. (2014), ApJ,
Binks & Jeffries (2014), MNRAS, 438, 11
Malo et al. (2014), ApJ, 792, 37
Mamajek & Bell (2014, in prep)